



Analysis of GLM Static Calibration Data

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Background: GLM



- **Lightning detection instrument built by Lockheed Martin for GOES-R**
 - **CCD-based detection**
 - **Advanced electronics for on-board processing of events**
 - **Ground processing algorithms for lightning detection**

<http://www.goes-r.gov>
Credit: Lockheed Martin

Instrument Parameter	Specification	References
ADC Resolution	14-bit System	
CCD Format	1372 rows x 1300 columns Variable-pitch array	Bredthauer, SPIE, v. 8453, 2012
Frame Rate	500 Hz	
Detection Wavelength	777.4 nm	Center of OI triplet



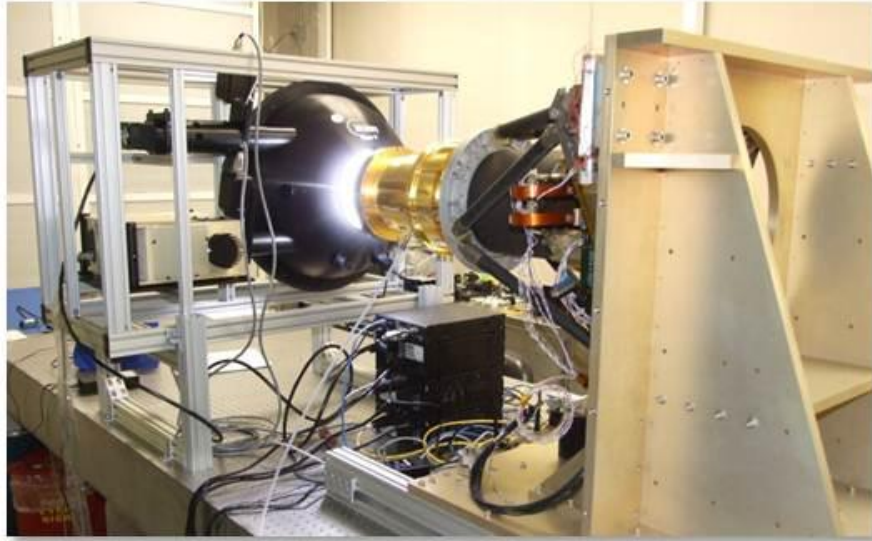
Calibration Overview and Additional Analysis



- **GLM Static Calibration Methodology**
 - Large sphere and Laser
- **Results of GLM testing**
 - Large sphere and laser
 - Lessons learned from supplementary measurements
- **Additional analysis of GLM vendor data**
- **Further Gain studies**
- **Updated plans for FM2**



GLM Vendor Instrument Calibration Methodology and Setup



GLM Sensor Unit during calibration testing

<http://www.goes-r.gov>

Credit: Lockheed Martin

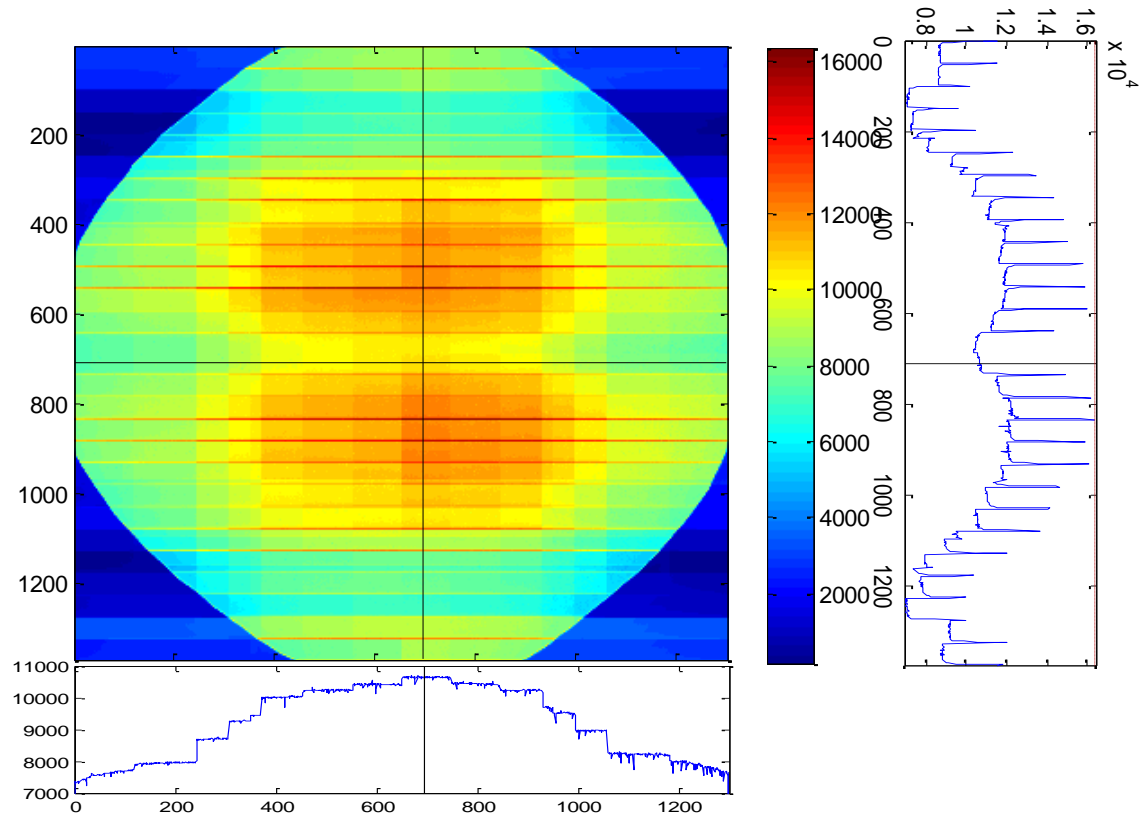
- **Employ large integrating sphere positioned to fill input optics**
 - Short sphere to optics separation distance
- **Measure average background images for range of radiances**
- **Calculate slope of count output versus reported radiance input**
 - Sphere monitoring photodiode reports input radiance to GLM
- **Replace Sphere by laser light delivery systems for gain test to refine sphere results**
 - Operate GLM as a power meter
- **Measure instrument counts versus input laser power in laser footprint**



Results of Large Integrating Sphere Illumination



- Largest pixels have greatest signal levels
 - Largest solid angle
- Bright horizontal lines observed
 - Vertical profile shows overshoot at 2nd pixel of each column in subarray
 - Signal amplitudes near ADC limit of 16383 for several rows



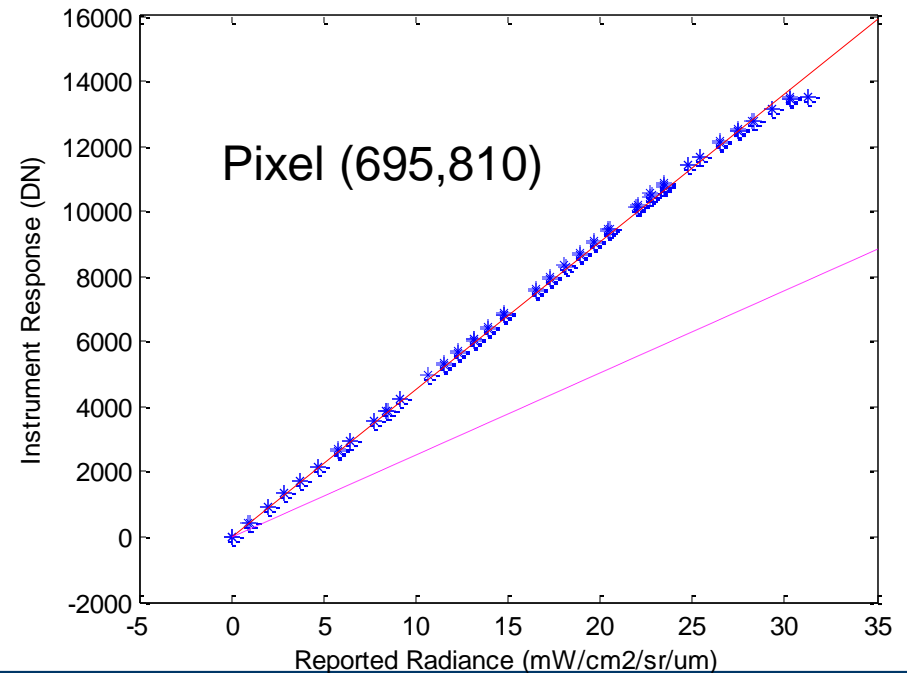
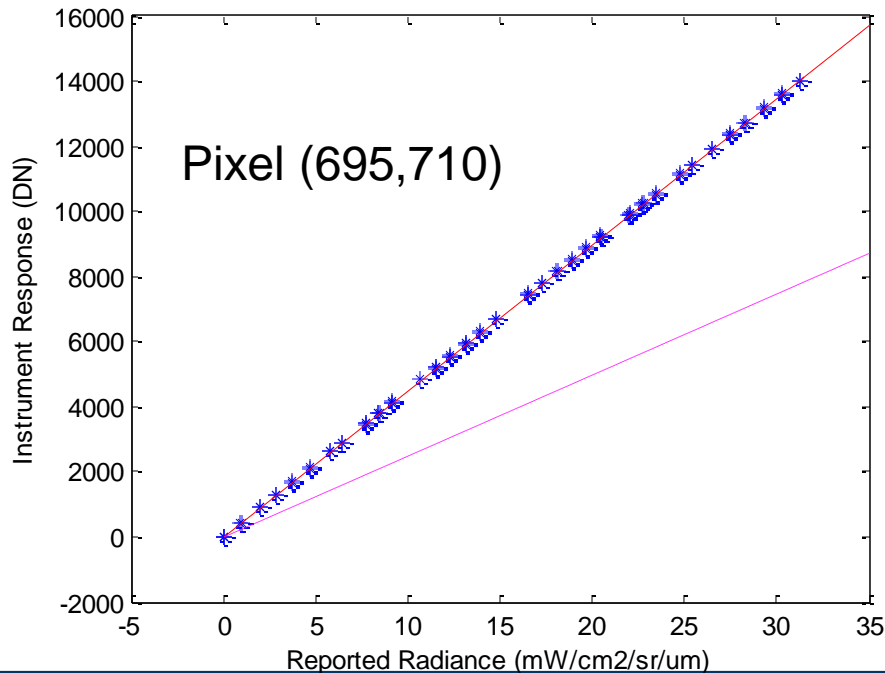
Setting near 20.5 mW/cm²/sr/μm



Analyzed Results: Instrument Response as a function of Reported Input Radiance



- **Response (in digital numbers) versus illumination level for two pixels**
 - **Left: Linear fit having no major deviations in response**
 - **Right: Linear fit having slight deviations at high illumination**
- **Calibration coefficient near 450 DN/(mW/cm²/sr/μm) for both pixels**
 - **Coefficients approximately 80% larger than expected**
 - **Dashed line shows expected instrument response**

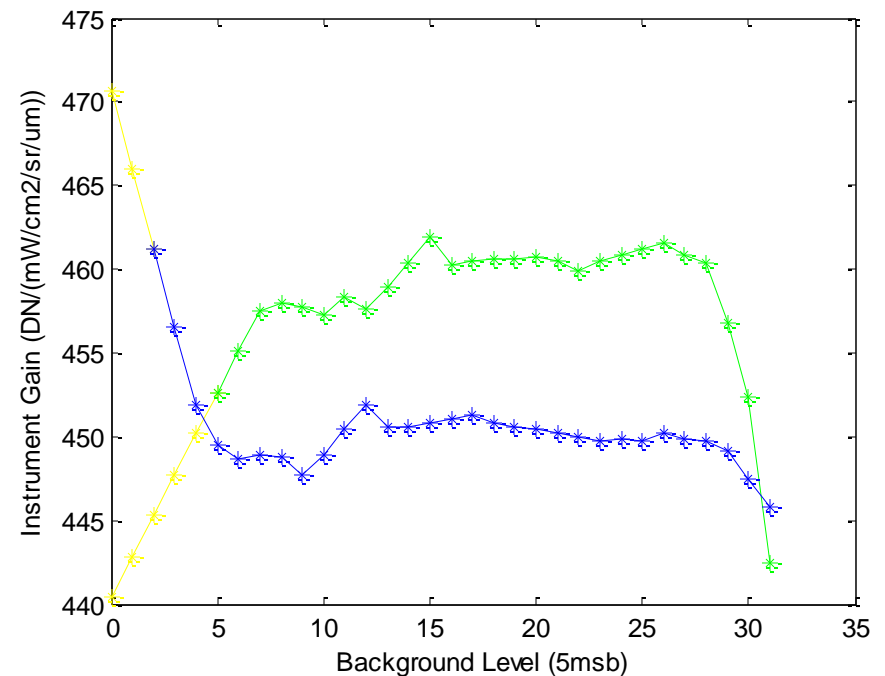
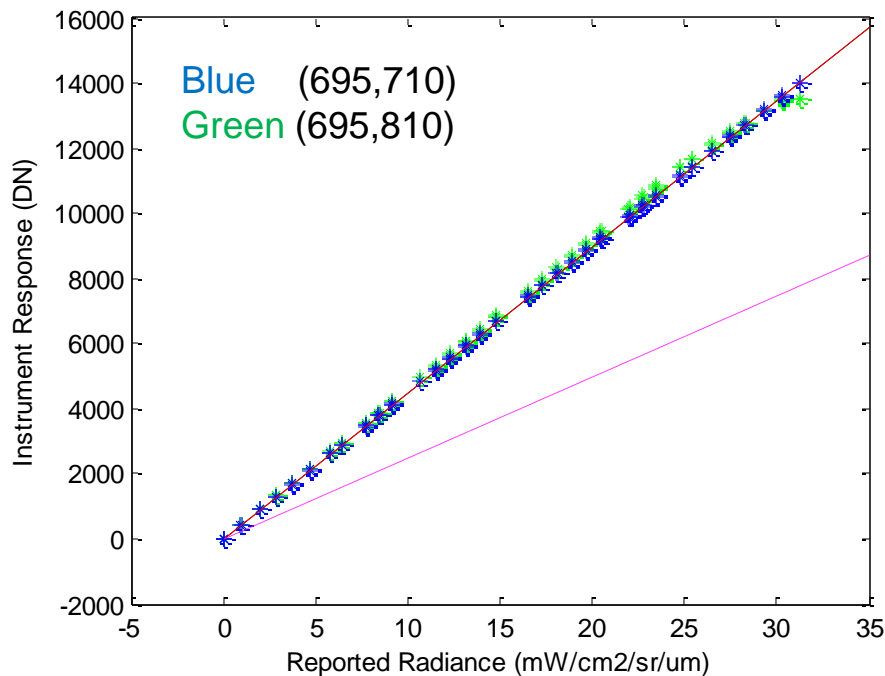




Investigated Data: Instrument Response as a function of Input Radiance



- **Left: Plotted response (in digital numbers) versus illumination level for two pixels**
- **Right: Plotted piecewise gain vs average background level**
 - Yellow points represent extrapolated values below the DC offset
 - Blue and green curves show the same pixels with the green curve has less headroom due to higher DC offset

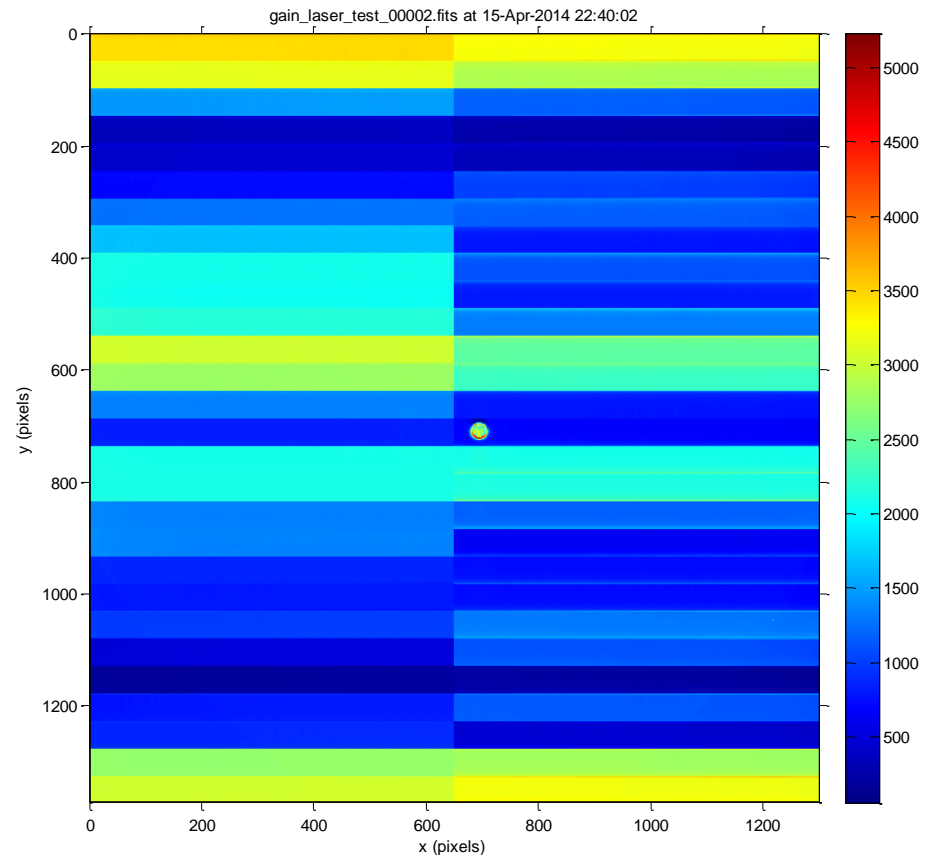




Added Laser Gain Measurement: Verify Instrument Transfer Value (e-/DN)



- **Simple analysis based on using GLM as a power meter**
- **Wavelength tuned to maximize transmission through spectral filters**
- **Only illuminates small number of pixels**
- **Must transfer gain to all pixels in CCD**

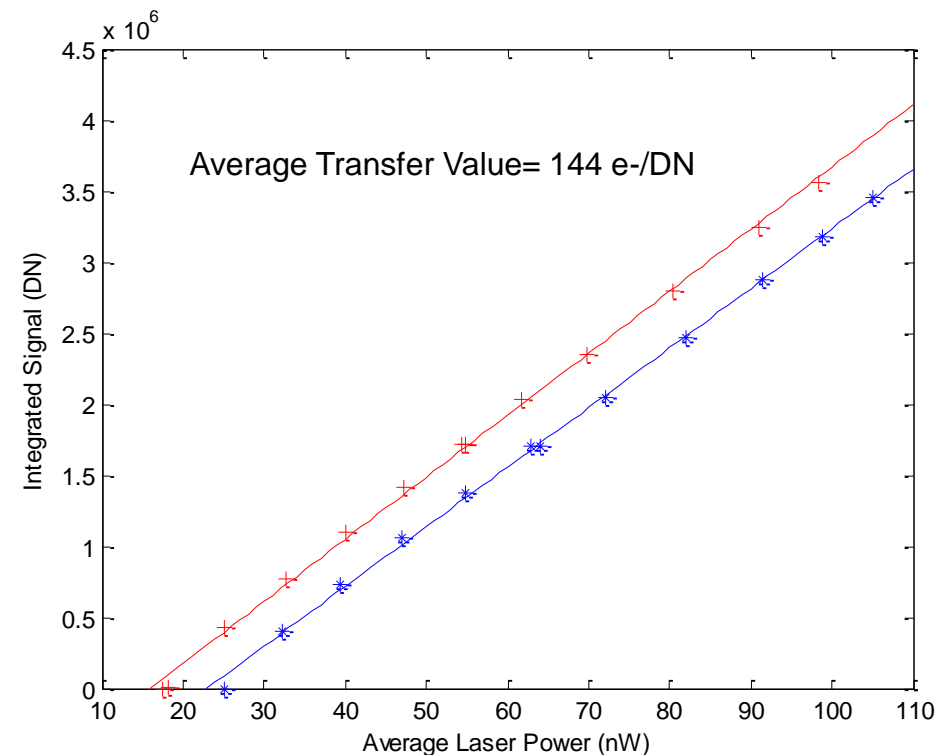
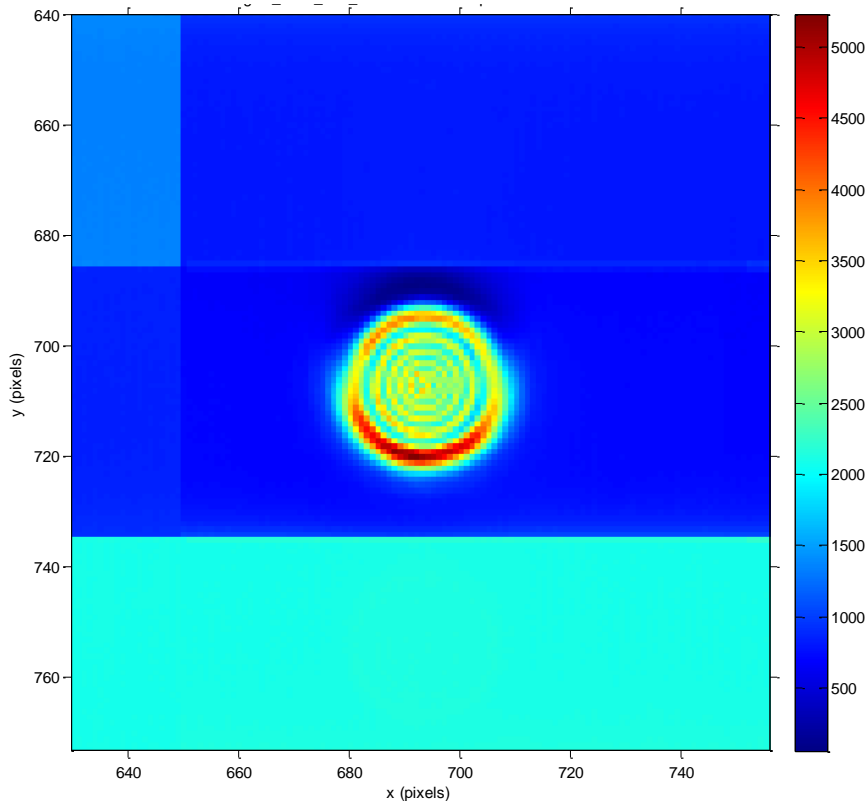




Laser Gain Measurements and Results

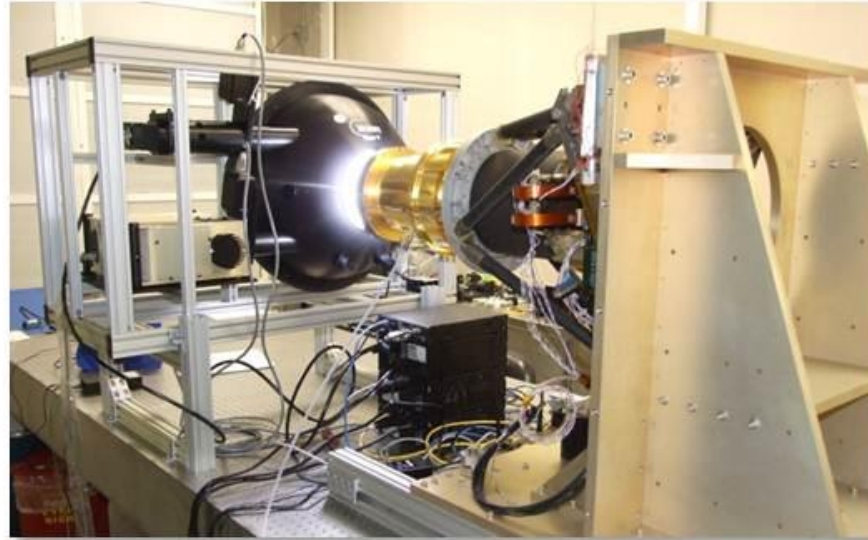


- Summed the response due to the laser light for each light power level
- Fitted data using linear equation ($y = a_1x + a_0$)
- Average transfer value for two measurement runs = 144 e-/DN
 - Sphere gain 1.8x of laser gain value





Lessons Learned from Supplementary Measurements



<http://www.goes-r.gov>
Credit: Lockheed Martin

- **Sphere sensor displayed correct values**
- **Differences between spectral filter on integrating sphere monitoring photodiode and instrument filters**
 - **Additional optical analysis was performed to verify**
- **Retro-reflections accounted for by monitoring photodiode**
- **Acceptance of out-of-band light by instrument**
- **GLM vendor reporting 80% correction due to test setup**

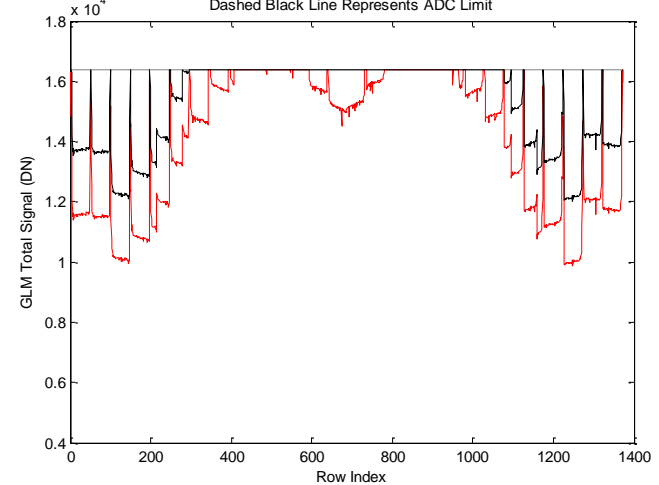


Best (left) and Worst (right) Extremes for Simulated Gain Cases and Impacts

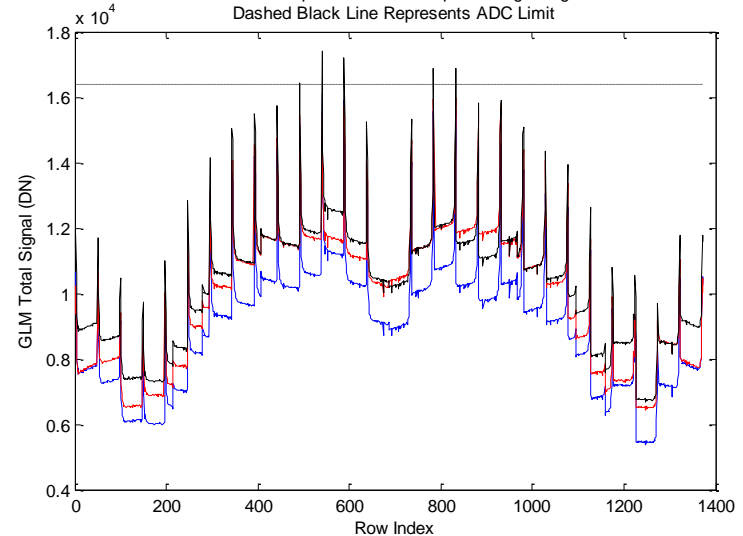


- **Correction by laser gain measurement increases input radiance by 1.8x**
- **Plots on right show unmodified dynamic range calculations with unmodified radiance values**
 - **GLM would saturate in brightest conditions**
 - **Some pixels would be blind under these conditions**
- **Corrected radiances (below) show GLM would not saturate with the radiance of a Lambertian source**

Maximum Signal Profile at 100% Albedo (Column Index = 692)
Redundant (Red) Side and Black Line Encompasses 90% of Expected Lightning Events
Dashed Black Line Represents ADC Limit



Maximum Signal Profile at 100% Albedo (Column Index = 692)
Primary (Blue) and Redundant (Red) Sides
Black Line Encompasses 90% of Expected Lightning Events
Dashed Black Line Represents ADC Limit





Summary



- **Updated static calibration methodology was executed by GLM vendor, accounting for details of test set up**
 - **Short separation distance between sphere and instrument**
 - Allowed light to retro-reflect from instrument back into sphere
 - Allowed light at high angles of incidence to enter instrument
 - **Gain discrepancy resolved by lab testing and analysis**
 - **Flight model 1 is well calibrated**
- **Static calibration plan was updated for FM2-4 from lessons learned**
 - **Increase separation distance between sphere and instrument**
 - **Perform laser gain measurements on each subarray of CCD**